

The Phases of Quantum Chromodynamics

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- ▶ Motivation:
 - ▶ High-temperature universe, i.e. fractions of a second after the Big Bang
 - ▶ High-density matter, i.e. in a neutron star
 - ▶ Understanding of QCD in extreme environments → deeper understanding of theory

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 - ▶ High-temperature universe, i.e. fractions of a second after the Big Bang
 - ▶ High-density matter, i.e. in a neutron star
 - ▶ Understanding of QCD in extreme environments → deeper understanding of theory
- ▶ Methods:
 - ▶ Almost no rigorous results
 - ▶ High temperatures, low densities: lattice calculations
 - ▶ High densities: analytic calculations

Finite temperature QFT

Short repetition of what has already been said for confinement/deconfinement:

- ▶ Partition function of a statistical system ($\beta = 1/T$):

$$Z = \sum_{\text{all states}} e^{-\beta E}$$

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- ▶ Rewrite using

$$e^{-\beta E} \rightarrow \langle \alpha | e^{-\beta H} | \alpha \rangle$$

and Euclidian time $\tau = it$

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- ▶ Final result:

$$Z = \int_{\phi(0)=\phi(\beta)} \mathcal{D}\phi \exp \left[- \int_0^\beta d\tau \int d^3x \mathcal{L}_E \right]$$

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Chemical potential

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- ▶ In canonical ensemble: particle number N kept constant

Chemical potential

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 - ▶ Creation of particle-antiparticle pairs; conserved quantity defined through number of particles minus number of antiparticles

- ▶ In canonical ensemble: particle number N kept constant
- ▶ In a field theory:
 - ▶ Creation of particle-antiparticle pairs; conserved quantity defined through number of particles minus number of antiparticles
- ▶ More convenient: allow particle number to fluctuate, but introduce weight factor similar to Gibbs factor for the energy

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Grand canonical ensemble

- ▶ Grand canonical ensemble: partition function defined by

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- ▶ μ is called the chemical potential associated with the charge N : energy change due to the introduction of particles to the system

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- ▶ Potentially many chemical potentials in a system; here, baryon chemical potential μ_B used, associated with baryon number N_B

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- ▶ μ is called the chemical potential associated with the charge N : energy change due to the introduction of particles to the system
- ▶ Potentially many chemical potentials in a system; here, baryon chemical potential μ_B used, associated with baryon number N_B
- ▶ Grand (canonical) potential = Helmholtz free energy:

$$\Omega(T, \mu) = -T \ln Z = -pV$$

- ▶ Minimized in equilibrium $\rightarrow p$ maximized

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Variables in the phase diagram

- ▶ Need to define driving parameters of the phase diagram
- ▶ Choose such that they are constant throughout the system even at phase coexistence and intensive
- ▶ One obvious variable: temperature T
- ▶ Density? No, because at phase coexistence, different density in different phases
- ▶ Chemical potential: connected to density, but constant across phase boundaries

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Chiral symmetry: conventions

- ▶ Fermionic field described by Dirac equation
- ▶ Weyl representation used:

$$\gamma^5 = \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix}, \psi(x) = \begin{pmatrix} \psi_L(x)' \\ \psi_R(x)' \end{pmatrix}$$

$$\psi_L(x) = \frac{1 - \gamma^5}{2} \psi(x)$$

$$\psi_R(x) = \frac{1 + \gamma^5}{2} \psi(x).$$

- ▶ ψ'_L and ψ'_R : left-handed and right-handed components

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Chiral symmetry: conserved currents

- ▶ Consider symmetry transformations:

$$U(1)_B : \psi(x) \rightarrow e^{i\alpha} \psi(x)$$
$$\psi(x) \rightarrow e^{i\alpha\gamma^5} \psi(x)$$

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- ▶ Associated currents:

$$\begin{aligned}j^\mu(x) &= \bar{\psi}(x)\gamma^\mu\psi(x) \\ j^{\mu 5}(x) &= \bar{\psi}(x)\gamma^\mu\gamma^5\psi(x)\end{aligned}$$

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- ▶ Divergences:

$$\begin{aligned}\partial_\mu j^\mu &= 0 \\ \partial_\mu j^{\mu 5} &= 2im\bar{\psi}\gamma^5\psi\end{aligned}$$

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- ▶ Divergences:

$$\begin{aligned}\partial_\mu j^\mu &= 0 \\ \partial_\mu j^{\mu 5} &= 2im\bar{\psi}\gamma^5\psi\end{aligned}$$

- ▶ The associated currents are conserved!

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Chiral symmetry: Left- and right-handed currents

- ▶ Define left- and right-handed currents

$$\begin{aligned}j_L^\mu &= \bar{\psi}_L \gamma^\mu \psi_L \\j_R^\mu &= \bar{\psi}_R \gamma^\mu \psi_R\end{aligned}$$

- ▶ We find:

$$\begin{aligned}j_L^\mu + j_R^\mu &= j^\mu \\ \partial_{\mu} j_L^\mu &= \partial_{\mu} j_R^\mu = 0.\end{aligned}$$

- ▶ Currents for left- and right-handed quarks are conserved separately!

Chiral symmetry in QCD

- ▶ Consider QCD with a doublet of massless quark flavours:

$$Q = \begin{pmatrix} u \\ d \end{pmatrix} = Q_L + Q_R, \quad Q_{L/R} = \frac{1 \mp \gamma^5}{2} \begin{pmatrix} u \\ d \end{pmatrix}$$

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- ▶ Transform separately under isospin transformations

$$U_L, U_R \in SU(2) : Q_L \rightarrow U_L Q_L, \quad Q_R \rightarrow U_R Q_R$$

- ▶ Chiral flavour symmetry of QCD: $SU(N_f)_L \times SU(N_f)_R$, N_f number of (massless) quark flavours

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$$U_L, U_R \in SU(2) : Q_L \rightarrow U_L Q_L, \quad Q_R \rightarrow U_R Q_R$$

- ▶ Chiral flavour symmetry of QCD: $SU(N_f)_L \times SU(N_f)_R$, N_f number of (massless) quark flavours
- ▶ Full symmetry group of QCD:

$$SU(3)_C \times SU(N_f)_L \times SU(N_f)_R \times U(1)_B$$

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Chiral symmetry breaking in QCD

- ▶ Strong attractive interactions between quarks and antiquarks, i.e. negative contribution to total energy
- ▶ Energy cost to create a pair of massless particles is very small

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Chiral symmetry breaking in QCD

- ▶ Strong attractive interactions between quarks and antiquarks, i.e. negative contribution to total energy
- ▶ Energy cost to create a pair of massless particles is very small
- ▶ Therefore, vacuum is populated by quark-antiquark pairs
- ▶ These have overall momentum and angular momentum of 0 \rightarrow they carry net helicity charge, pairs of left-handed quarks and left-handed antiquark, which is antiparticle of right-handed quark

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Chiral symmetry breaking in QCD 2

- ▶ Non-zero expectation value for quark-antiquark pairs:

$$\langle 0 | \bar{Q} Q | 0 \rangle = \langle 0 | \bar{Q}_L Q_R + \bar{Q}_R Q_L | 0 \rangle \neq 0$$

- ▶ Apply chiral flavour symmetries $U_L, U_R \in SU(2)$:

$$\langle 0 | \bar{Q}_L Q_R + \bar{Q}_R Q_L | 0 \rangle = \langle 0 | \bar{Q}_L U_L^\dagger U_R Q_R + \bar{Q}_R U_R^\dagger U_L Q_L | 0 \rangle$$

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- ▶ Apply chiral flavour symmetries $U_L, U_R \in SU(2)$:

$$\langle 0 | \bar{Q}_L Q_R + \bar{Q}_R Q_L | 0 \rangle = \langle 0 | \bar{Q}_L U_L^\dagger U_R Q_R + \bar{Q}_R U_R^\dagger U_L Q_L | 0 \rangle$$

- ▶ Only fulfilled for $U_L = U_R$!
- ▶ Appearance of condensate breaks $SU(N_f)_L \times SU(N_f)_R \rightarrow SU(N_f)_V$
- ▶ Spontaneous symmetry breaking: symmetry of Lagrangian not realized in ground state

▶ Assumptions:

- ▶ No long-range interactions, e.g. Coulomb forces \rightarrow true for QCD w/o Coulomb
- ▶ Lagrangian has a continuous, global symmetry
- ▶ Potential term selects a ground state (minimum potential)
- ▶ Ground state does not respect symmetry

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- ▶ Assumptions:
 - ▶ No long-range interactions, e.g. Coulomb forces \rightarrow true for QCD w/o Coulomb
 - ▶ Lagrangian has a continuous, global symmetry
 - ▶ Potential term selects a ground state (minimum potential)
 - ▶ Ground state does not respect symmetry
- ▶ Consequences (proof of classical, scalar case can be found in the report):
 - ▶ Occurrence of a massless bosonic particle called (Nambu-)Goldstone boson
 - ▶ True also for non-classical theories

Higgs mechanism

- ▶ Goldstone theorem: spontaneous breaking of a global continuous symmetry
- ▶ Higgs mechanism: spontaneous breaking of a local gauge symmetry!

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- ▶ Goldstone theorem: spontaneous breaking of a global continuous symmetry
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- ▶ Assumptions:
 - ▶ Lagrange function with local gauge symmetry
 - ▶ Ground state which is not invariant under gauge transformations

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- ▶ Goldstone theorem: spontaneous breaking of a global continuous symmetry
- ▶ Higgs mechanism: spontaneous breaking of a local gauge symmetry!
- ▶ Assumptions:
 - ▶ Lagrange function with local gauge symmetry
 - ▶ Ground state which is not invariant under gauge transformations
- ▶ Consequences:
 - ▶ New term in the Lagrangian: $\Delta\mathcal{L} = \frac{1}{2}m_A^2 A_\mu A^\mu$
 - ▶ Gauge bosons, described by A_μ , acquire mass!

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Simplified QCD

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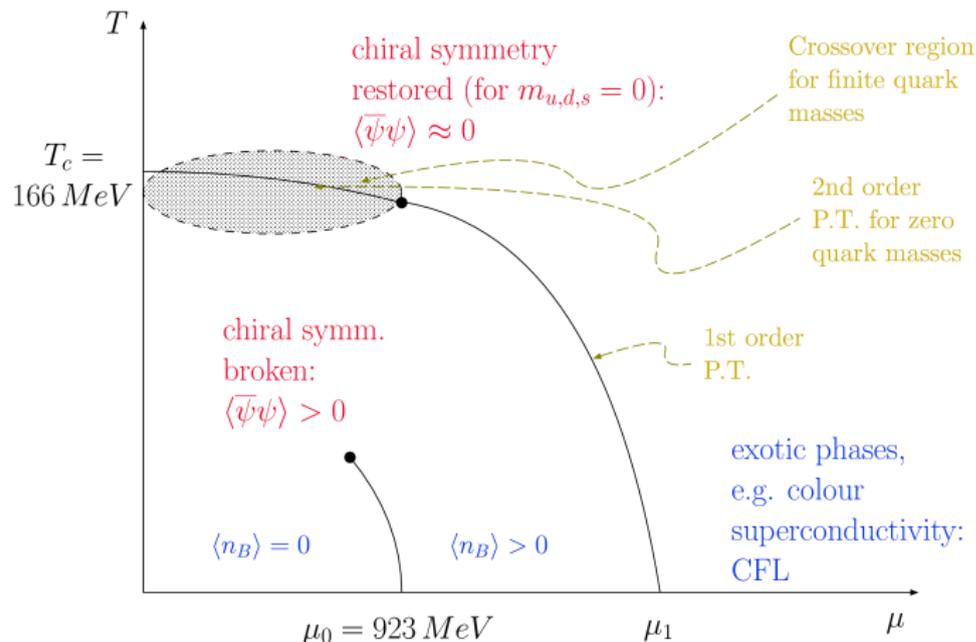
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- ▶ Electroweak interactions are ignored
- ▶ Two massless quarks u and d , no other quarks
- ▶ Leads to global $SU(2)_L \times SU(2)_R \times U(1)_B$ symmetry, broken down to $SU(2)_V \times U(1)_B$

The QCD Phase Diagram — schematically



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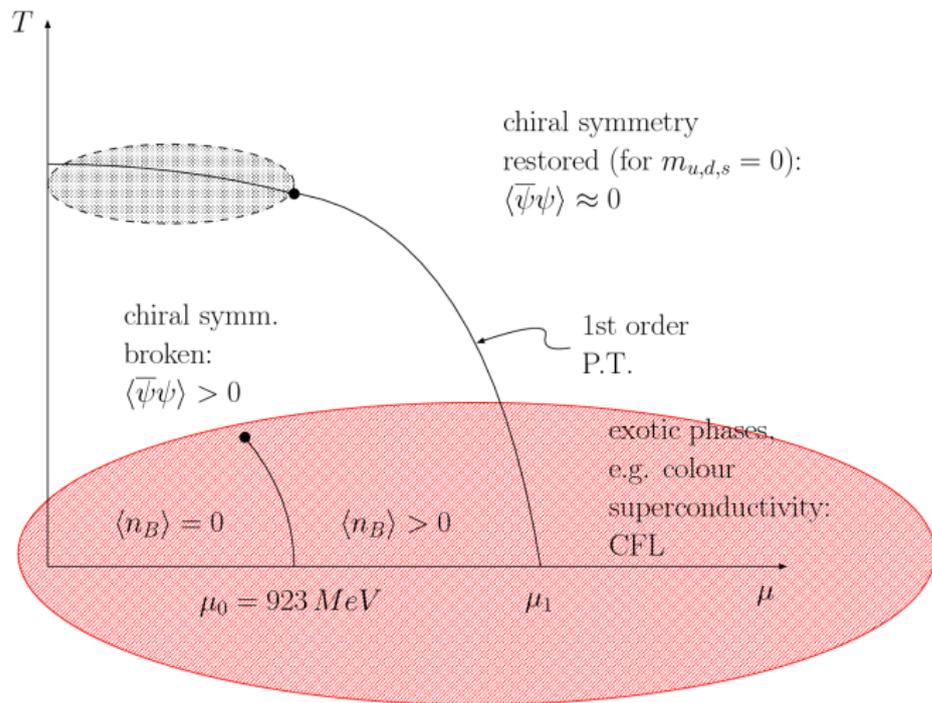
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Vacuum to nuclear matter: transition at μ_0

- ▶ Consider partition function:

$$Z = \sum_{\text{all states } \alpha} \exp\left(-\frac{E_\alpha - \mu N_\alpha}{T}\right)$$

- ▶ At $T = 0$: sum exponentially dominated by state which minimizes $E_\alpha - \mu N_\alpha$
- ▶ At $\mu = 0$: $N = E = 0$ with $n(\mu) = 0$

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- ▶ At $T = 0$: sum exponentially dominated by state which minimizes $E_\alpha - \mu N_\alpha$
- ▶ At $\mu = 0$: $N = E = 0$ with $n(\mu) = 0$
- ▶ $\mu > 0$:
 - ▶ $E_\alpha - \mu N_\alpha > 0$: still $N = E = 0$ dominating
 - ▶ $E_\alpha - \mu N_\alpha \leq 0$: other states contribute
- ▶ Therefore expect phase transition to $n(\mu) > 0$ at

$$\mu_0 := \min_{\alpha} \left(\frac{E_\alpha}{N_\alpha} \right)$$

- ▶ $n(\mu)$ is order parameter for the transition

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Vacuum to nuclear matter: value of μ_0

- ▶ Problem: find value of μ_0 for reduced and full QCD
- ▶ Reduced QCD:
 - ▶ Use $\frac{E}{N} = m_N - \frac{Nm_N - E}{N}$
 - ▶ Maximize second term $\epsilon = \frac{Nm_N - E}{N}$ which is binding energy per nucleon
 - ▶ Weizsaecker formula, w/o e.-m. interaction and for “infinitely large” nucleus: $\epsilon \approx 16$ MeV
 - ▶ Find first-order phase transition to $n_0 \approx 0.16 \text{ fm}^{-3}$ at

$$\mu_0 \approx m_N - 16 \text{ MeV} \approx 923 \text{ MeV}$$

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- ▶ Full QCD with Coulomb forces:
 - ▶ Infinite nucleus unstable due to Coulomb repulsion
 - ▶ Highest binding energy: iron nuclei \rightarrow adding electrons for neutrality, phase transition to iron solid at $\mu_0 \approx m_N - 8 \text{ MeV} \approx 931 \text{ MeV}$

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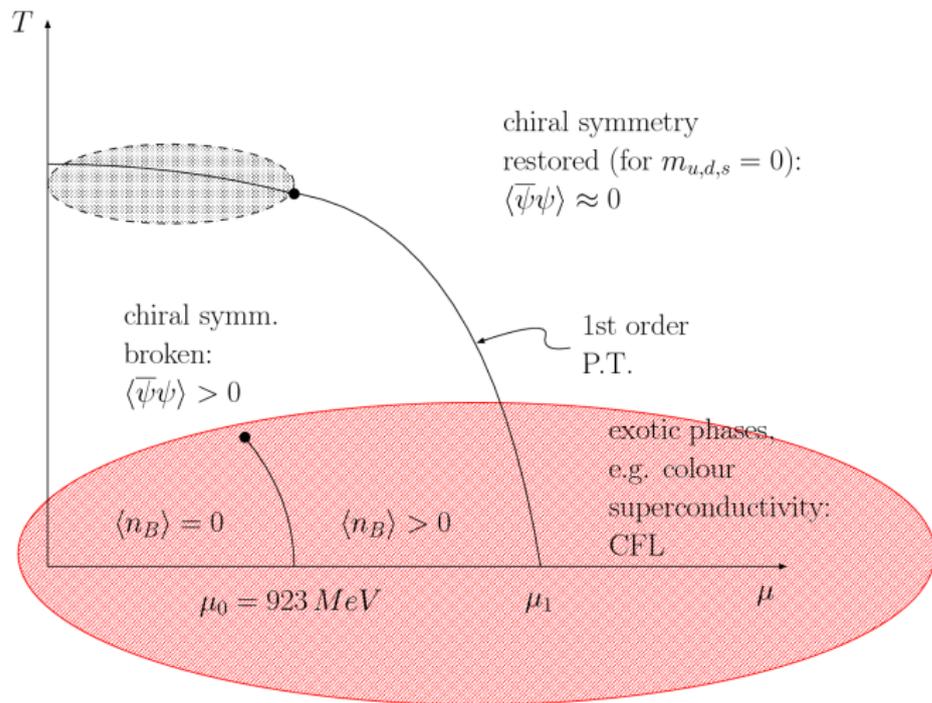
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- ▶ $\mu_0 < \mu < \mu_0 + 200$ MeV: very little known

- ▶ $\mu_0 < \mu < \mu_0 + 200$ MeV: very little known
- ▶ $\mu \gg \mu_0 + 200$ MeV: particles occupy high momentum states due to Fermi statistics
- ▶ High momentum \rightarrow asymptotic freedom \rightarrow chiral condensate vanishes
- ▶ Restoration of chiral symmetry (exact for massless quarks) is accompanied with phase transition at $\mu = \mu_1$
- ▶ Speculation: more phase transitions, exotic phases: colour superconductivity

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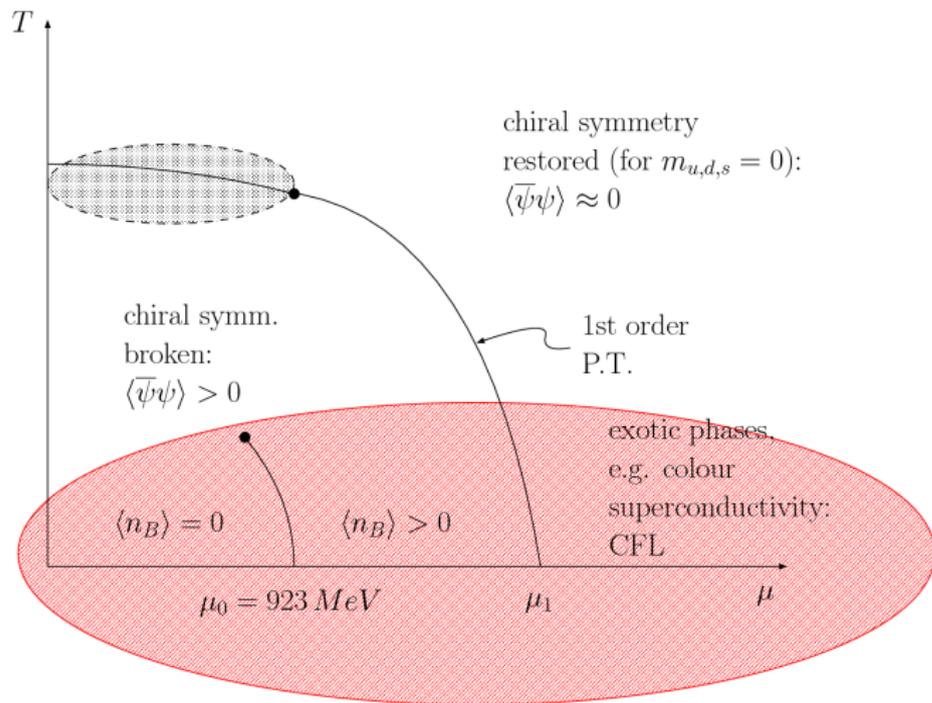
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The situation for low temperatures — Overview

- ▶ Asymptotic freedom for high momentum states still valid for finite temperatures
- ▶ Phase transition at $\mu = \mu_1$ not well understood, therefore little known about low-temperature behaviour

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- ▶ At $\mu = \mu_0$:
 - ▶ $n(\mu) > 0$ for finite T even for $\mu < \mu_0 \rightarrow$ no order parameter
 - ▶ But: discontinuities in 1st order phase transitions are assumed to appear as lines of 1st order p.t.s
 - ▶ First-order phase transitions should form lines terminated by critical point

The transition at $\mu = \mu_0$ for low temperatures

- ▶ Slope governed by Clausius-Clapeyron: $\frac{dT}{d\mu} = -\frac{\Delta n}{\Delta s}$

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The transition at $\mu = \mu_0$ for low temperatures

- ▶ Slope governed by Clausius-Clapeyron: $\frac{dT}{d\mu} = -\frac{\Delta n}{\Delta s}$
- ▶ In analogy to normal liquid-gas transition:
 - ▶ Nernst: $S \rightarrow 0$ for $T \rightarrow 0$
 - ▶ Therefore $\Delta s = 0$ at $T = 0 \rightarrow$ infinite slope

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 - ▶ Gas (high T) expected to have lower particle density $\rightarrow \Delta n < 0$
 - ▶ At 1st order transition, system absorbs heat: $\delta Q < 0 \rightarrow \Delta s < 0$
 - ▶ With $\frac{\Delta n}{\Delta s} > 0$, we find

$$\frac{dT}{d\mu} < 0$$

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- ▶ Expect line to terminate in critical point,
 $T_0 \approx \epsilon \approx 16$ MeV

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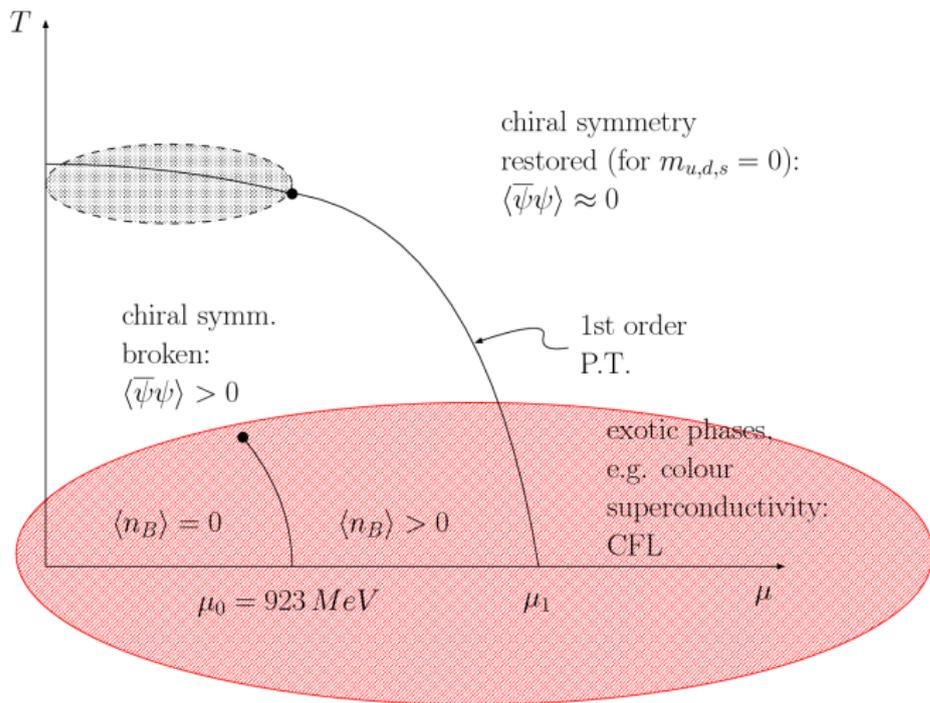
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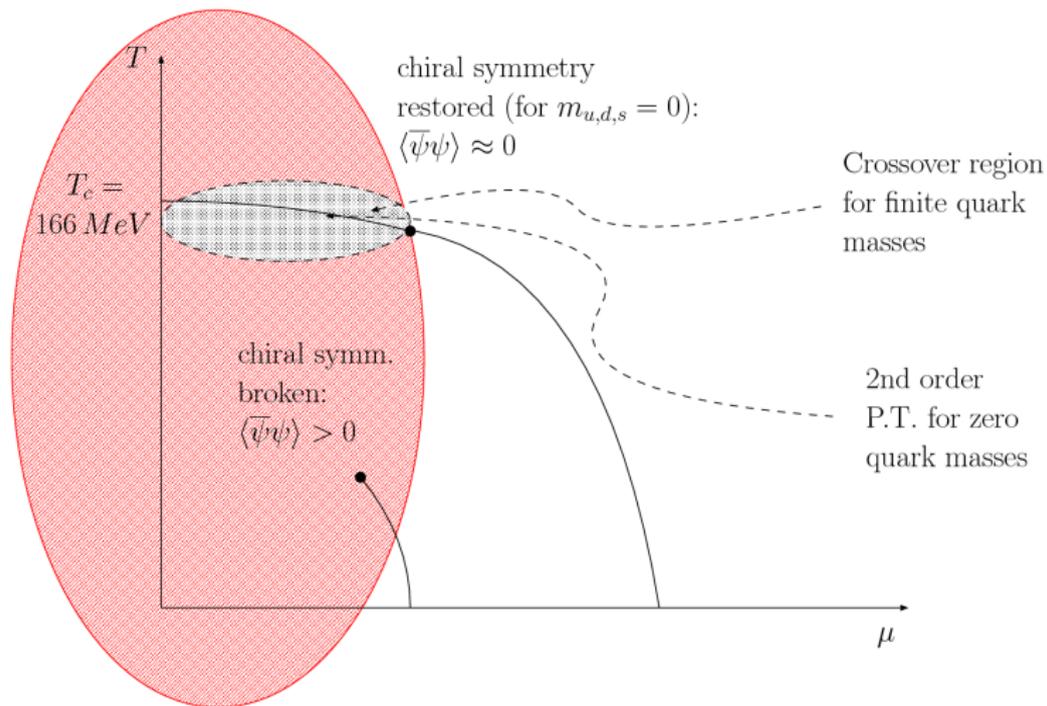
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- ▶ Raise temperature, keep chem. potential at $\mu = 0$
- ▶ Hadronic matter: dominated by pions as lightest mesons
- ▶ QGP: high-energy plasma of essentially free quarks and gluons, chiral symmetry restored

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- ▶ Assumptions
 - ▶ Ignore interactions in hadronic phase
 - ▶ Only pions formed in hadronic phase

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- ▶ Assumptions
 - ▶ Ignore interactions in hadronic phase
 - ▶ Only pions formed in hadronic phase
- ▶ Find transition: Helmholtz free energy $\Omega = -pV$ minimized
- ▶ Pressure is maximized \rightarrow calculate pressure in the two phases

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QGP transition: counting degrees of freedom

For non-interacting fields (n_f number of degrees of freedom):

Bosonic field $3P = \epsilon_B = n_f \frac{\pi^2}{30} T^4$

Fermionic field $3P = \epsilon_F = n_f \frac{7}{8} \frac{\pi^2}{30} T^4$

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► We find

$$T_c \approx 150 \text{ MeV}$$

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$$T_c \approx 150\text{MeV}$$

- ▶ First thought to be of first order
- ▶ Not seen experimentally!

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- ▶ We find

$$T_c \approx 150 \text{ MeV}$$

- ▶ First thought to be of first order
- ▶ Not seen experimentally!
- ▶ Current view:

Zero quark mass Second order

Non-zero quark mass No chiral symmetry \rightarrow no
symmetry breaking \rightarrow crossover region

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Summary of known transitions

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- ▶ Two lines of first order phase transitions along $T \approx 0$ and $\mu > 0$
- ▶ Second order p.t. (massless quarks) or crossover (finite quark masses) along $\mu \approx 0$ and $T > 0$
- ▶ Chiral condensate $\langle Q\bar{Q} \rangle$: vanishes in high- T - and high- μ -phases

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- ▶ Chiral condensate $\langle Q\bar{Q} \rangle$: vanishes in high- T - and high- μ -phases
- ▶ Chiral symmetry restoration in the region of high temperature and high chem. potential?

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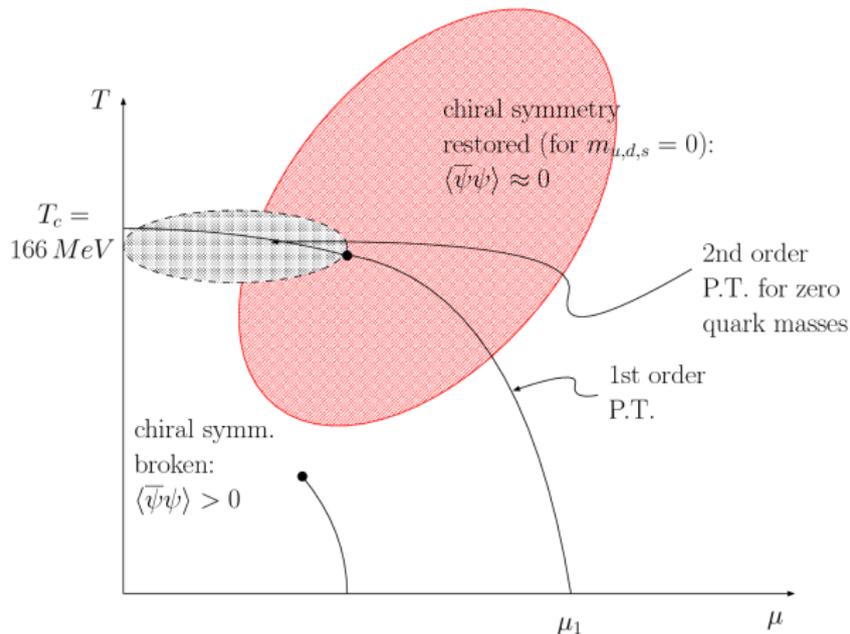
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Chiral symmetry restoration in the phase diagram

- ▶ Massless quarks:
 - ▶ Expect **one** line of phase transitions between regions of $\langle \overline{Q}Q \rangle = 0$ and $\langle \overline{Q}Q \rangle \neq 0$
 - ▶ Line of 2nd order p.t.s from the QGP transition and of 1st order p.t.s from the $\mu = \mu_1$ transition merge!
 - ▶ Meet in a tricritical point

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 - ▶ Line of 2nd order p.t.s from the QGP transition and of 1st order p.t.s from the $\mu = \mu_1$ transition merge!
 - ▶ Meet in a tricritical point
- ▶ Finite quark masses:
 - ▶ Chiral symmetry is never exactly restored
 - ▶ Line of first order phase transitions from $\mu = \mu_1$ does not terminate in tricritical point

The QCD Phase Diagram — schematically

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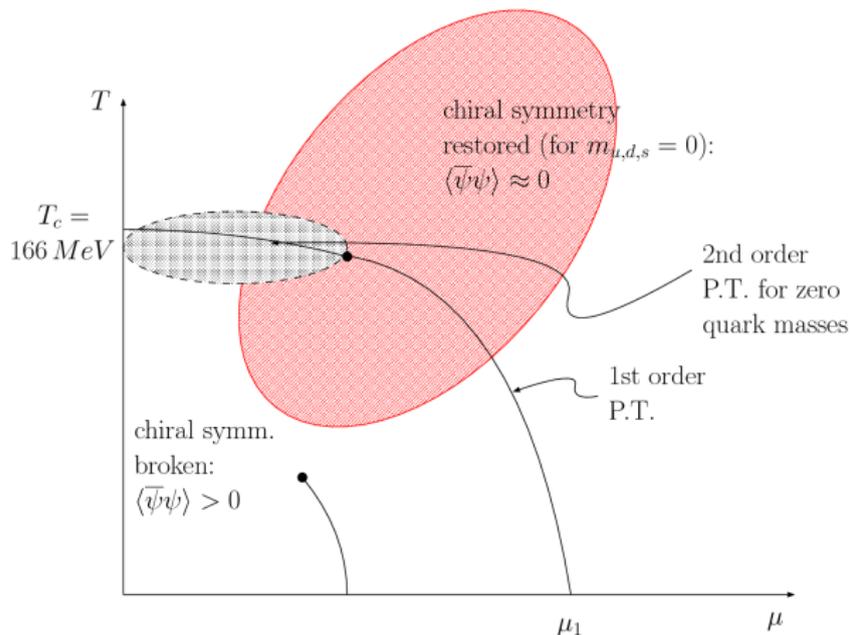
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Colour superconductivity

- ▶ Mechanism proposed for low temperatures and high chemical potentials: asymptotic freedom at high densities

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- ▶ Mechanism proposed for low temperatures and high chemical potentials: asymptotic freedom at high densities
- ▶ Two major types of colour superconductivity:
 - Two-flavour colour superconductivity 2SC Two massless quark flavours
 - Colour flavour locking Three massless quarks flavour. Interest due to chiral symmetry breaking by other mechanism than chiral condensate.

Colour superconductivity

- ▶ Mechanism proposed for low temperatures and high chemical potentials: asymptotic freedom at high densities
- ▶ Two major types of colour superconductivity:
 - Two-flavour colour superconductivity 2SC Two massless quark flavours
 - Colour flavour locking Three massless quarks flavour. Interest due to chiral symmetry breaking by other mechanism than chiral condensate.
- ▶ Densities difficult to produce in earthly laboratories
- ▶ Only due to gravitational collapse: neutron star

Ordinary superconductivity

- ▶ Ordinary solid: ideal, non-interacting Fermi gas approximation becomes valid for small temperatures
- ▶ For some materials: at T_c , weak attractive force (mediated by ion lattice) leads to formation of condensate of Cooper pairs
- ▶ Described by BCS theory for type I superconductivity
- ▶ Physical effects:
 - ▶ Gap in the excitation spectrum: binding energy of Cooper pairs
 - ▶ Zero electrical resistance
 - ▶ Meissner-Ochsenfeld effect: expulsion of magnetic fields

Superconductivity in the QCD case

- ▶ Limit of high density: asymptotic freedom, expect free Fermi gas

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Superconductivity in the QCD case

- ▶ Limit of high density: asymptotic freedom, expect free Fermi gas
- ▶ Attractive forces lead to formation of quark pair condensate:
 - ▶ Perturbative effects have small contribution
 - ▶ Non-perturbative (instanton) effects, proposed in late 1990s, have very large effect

Superconductivity in the QCD case

- ▶ Limit of high density: asymptotic freedom, expect free Fermi gas
- ▶ Attractive forces lead to formation of quark pair condensate:
 - ▶ Perturbative effects have small contribution
 - ▶ Non-perturbative (instanton) effects, proposed in late 1990s, have very large effect
- ▶ Colour charge superconductivity and Meissner-Ochsenfeld are not observable
- ▶ Energy gap: $\Delta = \frac{E}{2}$, E binding energy
- ▶ Estimate: $\Delta \approx 100$ MeV

Colour flavour locking

- ▶ Single-gluon exchange and non-perturbative effect provide attractive force between quarks
- ▶ Leads to formation of a condensate (analogous to Cooper pairs; (α, β) refer to colour, (i, j) to flavour):

$$\langle \psi_{iL}^{a\alpha}(\vec{p}) \psi_{jL}^{b\beta}(-\vec{p}) \epsilon_{ab} \rangle = -\langle \psi_{iR}^{a\alpha}(\vec{p}) \psi_{jR}^{b\beta}(-\vec{p}) \epsilon_{ab} \rangle \propto \epsilon^{\alpha\beta A} \epsilon_{ijA}$$

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- ▶ Use

$$\epsilon^{\alpha\beta A} \epsilon_{ijA} = \delta_i^\alpha \delta_j^\beta - \delta_j^\alpha \delta_i^\beta$$

- ▶ Locking of flavour and colour indices!

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- ▶ Locking of flavour and colour indices!
- ▶ Breaking of symmetries:

$$SU(3)_c \times SU(3)_L \times SU(3)_R \rightarrow SU(3)_{\text{color,L,R}}$$

- ▶ Chiral symmetry is broken, but by another mechanism

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- ▶ Spontaneous symmetry breaking:
 - ▶ Global symmetry: massless Goldstone boson
 - ▶ Local gauge symmetry: gauge bosons acquire mass

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 - ▶ Colour symmetry broken \rightarrow gluons acquire mass
 - ▶ Chiral symmetry broken: an octet of Goldstone bosons, oscillations of the diquark condensate
- ▶ If quark masses are non-zero, pseudo-Goldstone bosons are created with finite, but small, masses
- ▶ Can be considered as physical mesons with finite, computable masses: at $\mu = 400$ MeV,
 $m_{K^\pm} \approx 5 \dots 20$ MeV

Neutron star cooling

- ▶ Neutron star:
 - ▶ Star between 1.44x and 3x mass of the sun
 - ▶ Density $\approx 10^{12} \text{kg/cm}^3$ in the core
 - ▶ Radius 10 km

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- ▶ Neutron star cooling can be observed, e.g. fitting thermal radiation spectrum to blackbody radiation
- ▶ Takes place via neutrino emission
- ▶ Rate of cooling depends on heat capacity

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- ▶ Neutron star cooling can be observed, e.g. fitting thermal radiation spectrum to blackbody radiation
- ▶ Takes place via neutrino emission
- ▶ Rate of cooling depends on heat capacity
- ▶ In CFL: all quarks acquire gap with $\Delta \gg T$, leading to low thermal excitation
- ▶ No contribution to specific heat from quarks, specific heat dominated by electrons and pseudo-Goldstone bosons

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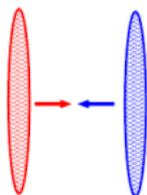
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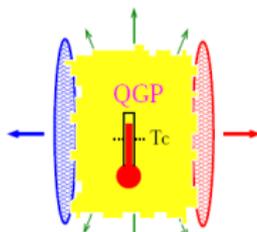
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- ▶ Objective: study QGP at high temperatures and low chemical potentials
- ▶ Collide heavy nuclei (sulphur, lead, gold) with high energies (A : nucleus weight):
 - ▶ Old: SPS @ CERN: CM energy $2A \dots 18 A$ GeV
 - ▶ Current: RHIC @ Brookhaven, NY: $200A$ GeV
 - ▶ Future: LHC (ALICE) @ CERN: $5500A$ GeV
- ▶ Other experiments will at some point attempt to study high-density matter, i.e. CBM/FAIR at GSI, Darmstadt

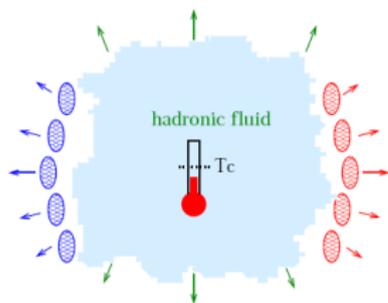
Stages of a RHIC collision



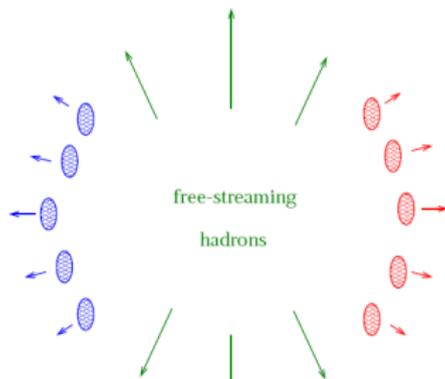
(a)



(b)



(c)



(d)

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Thank you for your attention